

2013.2.21

Study of the Venus cloud upper haze 金星雲・上部もや層の研究

Seiko Takagi, A. Mahieux, S. Robert, V. Wilquet,
R. Drummond, A.C. Vandaele, N.Iwagami

1. The University of Tokyo
2. Belgian Institute for Aeronomy

Early observation of the Venus cloud upper haze

- O'leary et al., 1975—Mariner10
 - Mariner10 found out about upper haze layer.
- Travis et al., 1979—Pioneer Venus polarization observation
 - Refractive index: 1.44, particle size: 0.4 μm
- Kawabata et al., 1980—Pioneer Venus polarization observation (UV)
 - Refractive index: 1.45 ± 0.04 (\rightarrow sulfuric acid), particle size: 0.23 μm , optical thickness: polar region 1°, equatorial region: 0.2
- Mukai et al., 1981—Comparison of observation and calculation
 - Haze is not sulfuric acid?
- Crisp., 1986—Comparison of observation and calculation (UV)
 - Optical thickness: 1–12
- Pollack et al., 1993—updated Crisp. (1986)
 - Optical thickness: 10

Recent observation of the Venus cloud upper haze

Wilquet et al., 2009/2012

- Venus Express
- Solar Occultation at InfraRed(SOIR)
- Upper haze: vertical and latitude distribution of extinction, particle size

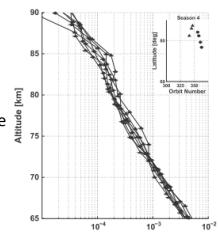
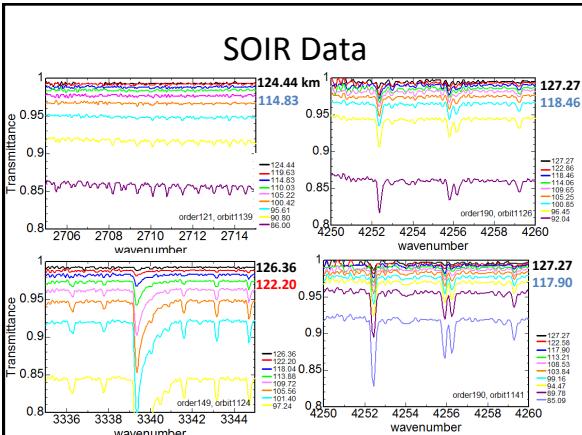
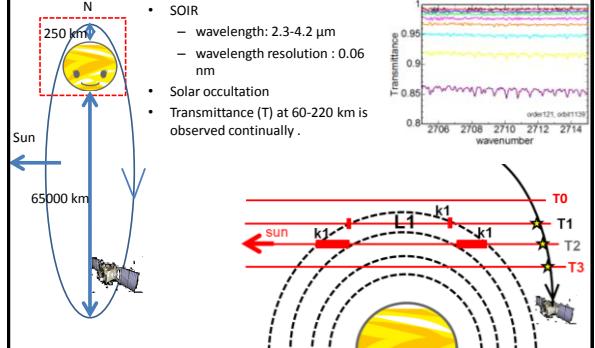


fig. Example of vertical distribution of upper haze

Venus Express Solar Occultation at InfraRed (SOIR)



Data analysis

