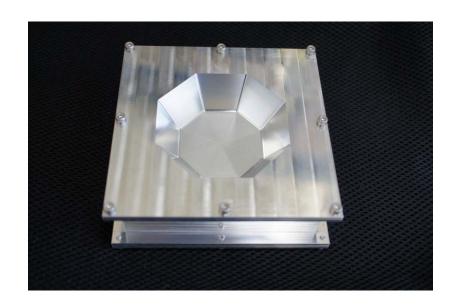
### Development of an energetic electron detector with 2-pi steradian field-of-view for planetary explorations by S. Kasahara

#### **ABSTRACT**

Generation of energetic electrons (up to > MeV) in planetary magnetospheres is a longstanding mystery, since their seed energy is of the order of 100 eV (solar wind) or much less (ionosphere). Furthermore, such high-energy electrons have significant influences for ionospheric dynamics as well as the long-term development of planets'/moons' surface and atmosphere. Since velocity distributions of electrons are not omnidirectional in many cases, observed velocity distribution function is the important information which provides specific restriction on the generation mechanisms and effects on planets and moons. In most cases, however, spacecraft for planetary explorations are 3-axis stabilised, and therefore it is difficult for traditional telescopetype electron detectors to obtain broad field-of-view (FOV). Toward future planetary explorations, here we design and fabricate a novel electron detector (20 - 100 keV) which can cover almost 2-pi steradian without the spacecraft spin. We develop this instrument in a step manner. First we will test the performance of this detector in a sounding rocket experiment on pulsating aurora, and then plan to downsize by using ASIC, and also extend the measurement energy up to ~ MeV by stacking detectors.

# Development of an energetic electron detector with 2-pi steradian field-of-view for planetary explorations

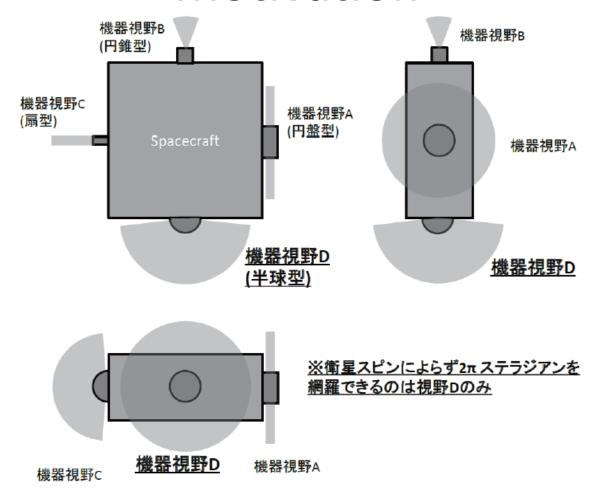
S. Kasahara (ISAS/JAXA)



#### **Points**

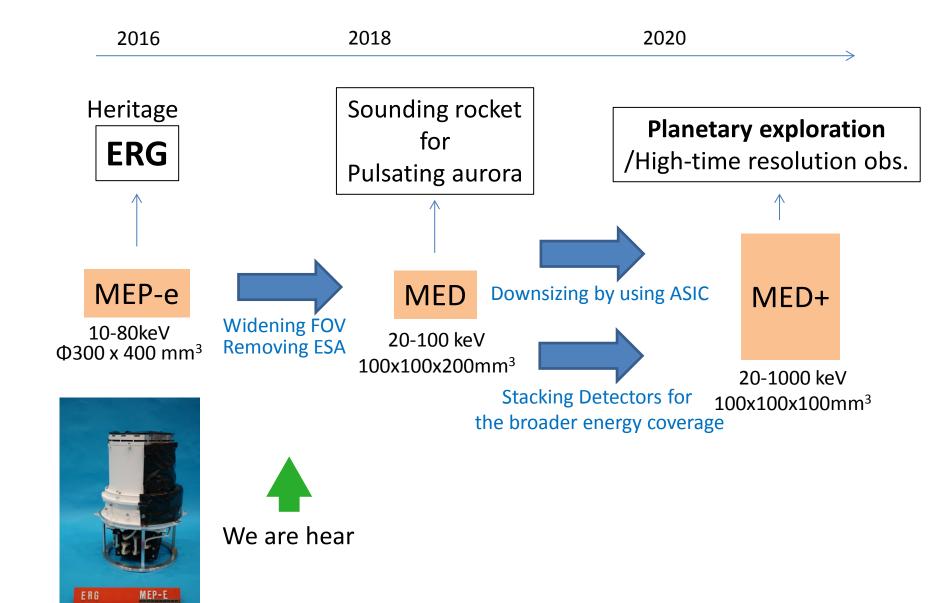
- Spacecraft for planetary explorations are mostly 3axis stabilised
- Wide field-of-view under the non-spin platform is thus required for particle instruments
  - Note that such instruments are also important for hightime resolution observations even in the case of spinstabilised spacecraft
- We develop the energetic particle instrument with 2-pi str field-of-view

#### **Motivation**

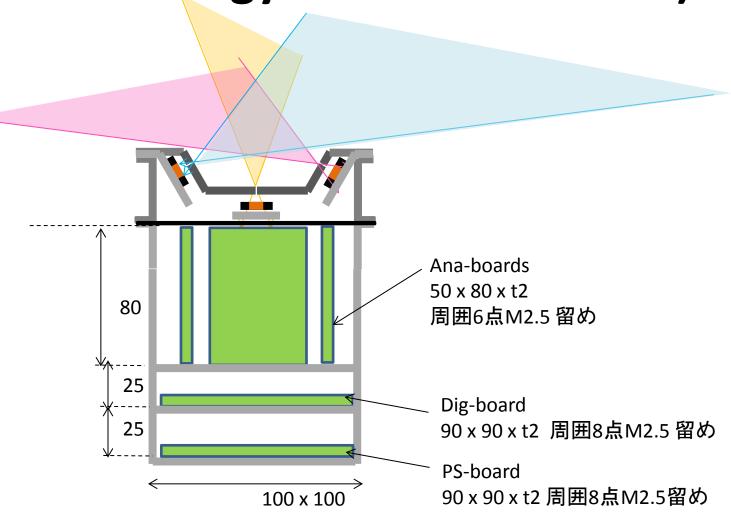


- Previous particle instruments cover only a small part of full solid angle if the platform does not spin
- Since spacecraft do not spin in most planetary missions, wide field-of-view is essential for particle sensors

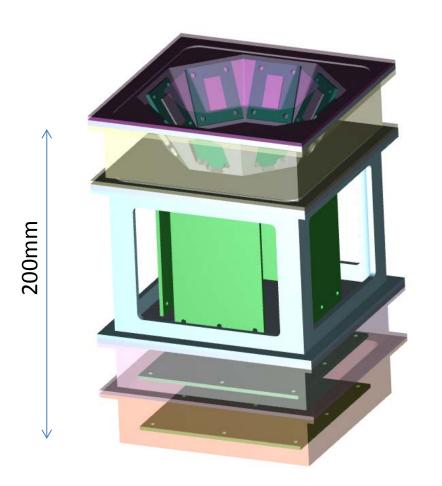
#### **Strategy**

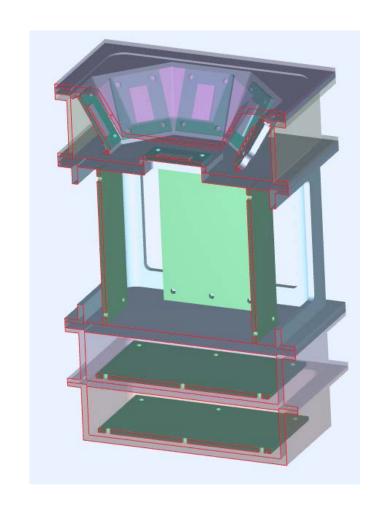


## Schematic view of MED (medium-energy electron detector)

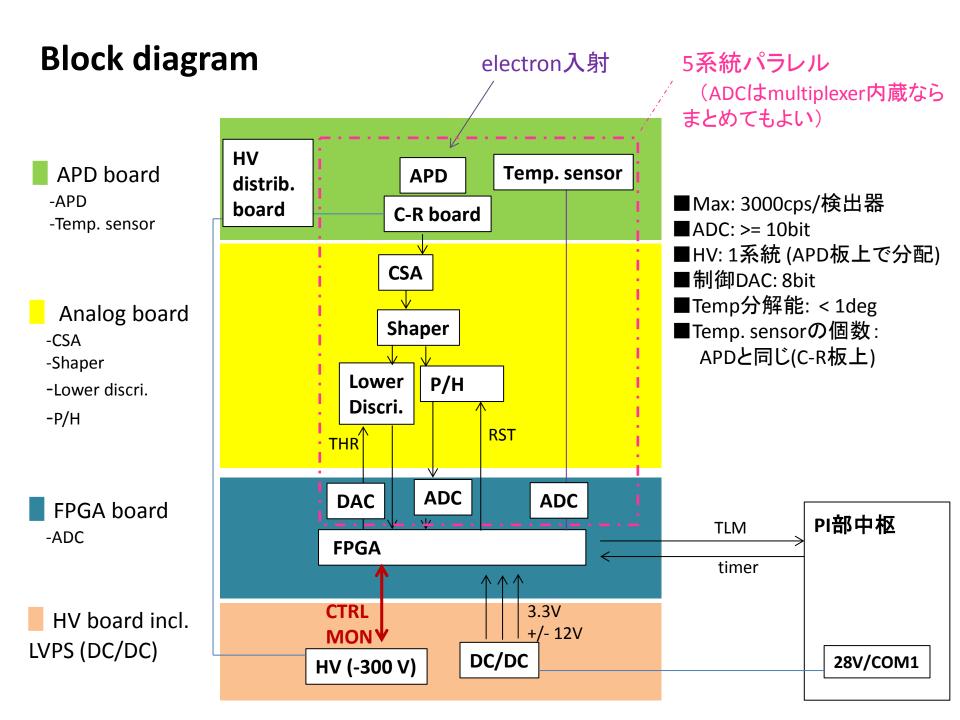


#### Sensor structure

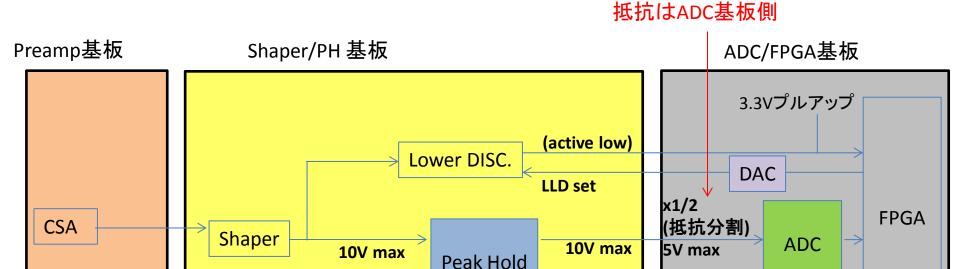




Aperture optics fabrication ongoing



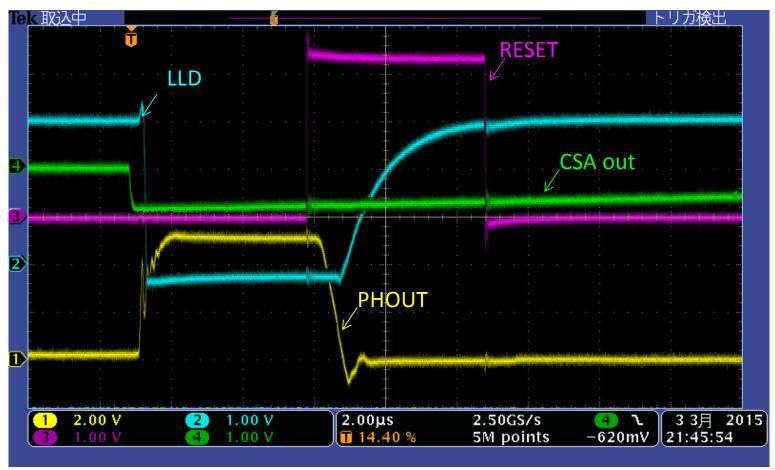
#### **Analogue circuit**



(1素子あたり8ch)

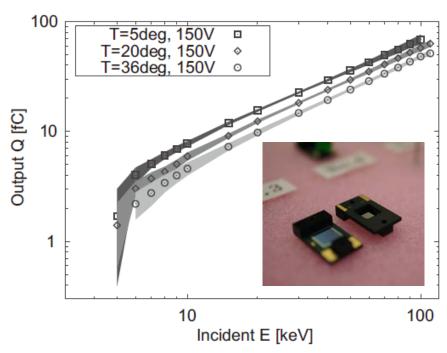
PH reset (3.3 V)

#### Signal time line (MEPe/ERG)

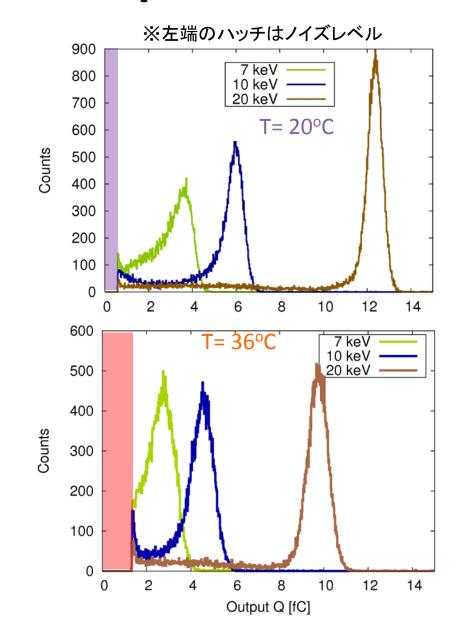


- The electron energy is converted into the voltage at CSA
- The CSA output is shaped and its peak value is held at Peak-holder(PH)
  - Thus the PH output value corresponds to the electron energy
- The digital electronics reads the PHOUT value when UD "LOW" is detected
  - Then send the RESET signal to flush the PHOU

#### **Detector: Avalanche photodiode**



- リーチスルータイプAPDを用いる
- 印加 HV 150-250V
- 面積~5 x 5 mm²
- 空乏層の厚み70um
- 不感層の厚み~0.2um/2um
- Energy resolution
  - < 20% (>20 keV)
  - < 30% (10-20 keV)</p>



#### **Specification**

Items		
Target	Electron	
Energy	20 keV – 100 keV	
coverage		
Energy	10-20%	depending on the
resolution		incident energy
Sensitivity	~ 10 <sup>-3</sup> cm <sup>2</sup> -sr	Per detector
FOV	80 deg cone	Centre of FOV points the
	(envelope)	direction perpendicular to
		the rocket spin axis.
Size	100mm x 100mm x	
	200mm	
Weight	3kg	
Power	8W	
Data rate	300kbps	

#### Summary

- We have designed energetic electron detector with 2-pi str FOV, which should be installed on future planetary explorers
- We're preparing for
  - Tests of analogue circuit boards
  - Check of APD performance
    - Especially the EUV rejection property