Reconsideration for causalities of occurrence features of Io-related Jupiter's radio emission

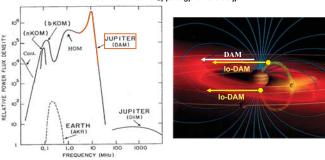
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The following questions; 'What kind of magneto-ionic wave Jupiter's auroral radio emission is?' and 'How the radio emission is generated?' have been long years of subjects. I have investigated the subjects based on numerical calculations using several kinds of magnetic field and plasma density models, however, the questions have not been resolved yet: a hypothesis of a special energy transporter which does not meet with the observation results was needed. Recently Jupiter's new magnetic field model 'JRM09' was proposed based on the JUNO Jupiter explorer conducting in-situ magnetic field measurements near Jupiter (Connerney+, GRL, 2018). We have tried to make a 3D raytracing analysis for Io-DAM using the JRM09 model. The preliminary analyses show that R-X mode waves are preferable as Io-DAM and the JRM09 model gives more natural explanations for the origin of Io-DAM, though there still remain some questions on restriction of 'Io-DAM' and on origin of Io-C; i.e., some additional energy input process(es) so as to meet the ray emitting conditions with the observed Io-DAM sources.

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Reconsideration for causalities of occurrence features of Jupiter's lo-related radio emission

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Io-DAM: Unexplained Characteristics

Occurrence probability · · · Localized Sources

Fig. Spectral profile of Jupiter's radio emission (Kaiser, 1993)

CML 1965 (deg)

270

90

(deg)

Purpose of this study

ABSTRACT

The following questions; 'What kind of magneto-ionic wave

Jupiter's auroral radio emission is?' and 'How the radio emission is generated?' have been long years of subjects. I have investigated

the subjects based on numerical calculations using several kinds of

magnetic field and plasma density models, however, the questions have not been resolved yet: a hypothesis of a special energy transporter which does not meet with the observation results was needed. Recently Jupiter's new magnetic field model 'JRM09' was

proposed based on the JUNO Jupiter explorer conducting in-situ magnetic field measurements near Jupiter (Connerney+, GRL, 2018). We have tried to make a 3D raytracing analysis for Io-DAM using the JRM09 model. The preliminary analyses show that R-X mode waves

are preferable as Io-DAM and the JRM09 model gives more natural explanations for the origin of Io-DAM, though there still remain some

questions on restriction of 'lo-DAM' and on origin of lo-C.

Origin of lo-DAM · · · No reliable direct obs.

- Investigation of origin (generation & propagation processes) of lo-DAM using numerical analyses ("ray trace", "polarization trace")
 - Magneto-ionic mode ? (R-X or L-O?)
 - Generation conditions? (source location, initial ray direction)
 - Plasma conditions in source & propagation regions? ···→ next step (polarization analysis)

Previous numerical analysis for Io-DAM

Method: 3D Ray & Polarization tracing (Misawa+, 2008):

- •model of magnetic field : VIP4 (Connerney+ 1998)
- •model of plasma density : Ne_{max}=6x10⁵[cm⁻³], H~960km (Eshleman+ 1979)
- wave mode : R-X & L-O

Results:

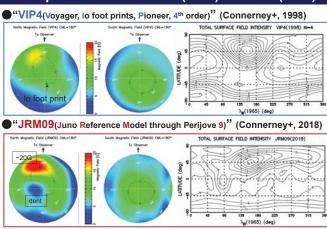
- •wave mode : "R-X"
- •wave direction : ~90° to B
- •lead angle : ~20° ~

•special requirements : "selective energy input" for the longitudes of northern lo-DAM sources

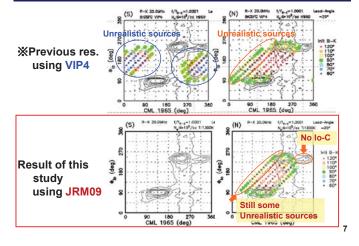
Comparison between VIP4(1998) & JRM09(2018)

Fig. Occurrence prob. for DAM (obs. by Tohoku Univ. in 1974-1994)

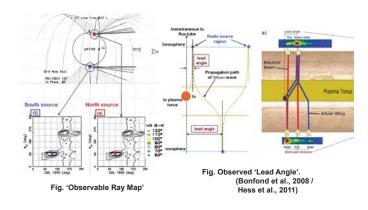
Polarization · · · RH elliptical pol.



Results for mag. models of VIP4 & JRM09



Explanation of 'Observable Ray Map' & 'Lead Angle'



Summary

- Purpose: Investigation of origin (source location, wave mode, generation & propagation conditions) of Jupiter's Io-DAM emission
- Method: 3D raytracing

Conditions: •R-X & L-O mode waves from S & N hemis.

- Including Lead-angle
- ·Emission to all-directions
- •Diffusive equilibrium Ne model
- New mag. Model "JRM09"
- Result: Io-DAM is generated:
 - •R-X mode from N-hemisphere
 - ·Lead-angle ~20° ~
 - •Cone half angle ~90° ~
 - •'JRM09' like mag.-field. is more plausible. (•in tenuous Ne ionosphere (by Pol.-analysis))

However, there are still remaining questions:

- •Some additional conditions to restrict 'lo-DAM source'?
- •More precise mag.-field model may be further needed?

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